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## Photoelectron spectroscopy of HCCS radical

M. Drissi ; G. A. Garcia ; B. Gans ; S. Boyé-Péronne ; H. R. Hrodmarsson ; L. Nahon ; J.-C. Loison 



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M. Drissi,<sup>1,a)</sup> G. A. Garcia,<sup>1</sup> B. Gans,<sup>2</sup> S. Boyé-Péronne,<sup>2</sup> H. R. Hrodmarsson,<sup>3</sup>  
L. Nahon,<sup>1</sup> and J.-C. Loison<sup>4,b)</sup>

## AFFILIATIONS

<sup>1</sup>Synchrotron SOLEIL, L'Orme des Merisiers, Départementale 128, 91190 Saint-Aubin, France

<sup>2</sup>Institut des Sciences Moléculaires d'Orsay (ISMO), CNRS, Université Paris-Saclay, 91405 Orsay, France

<sup>3</sup>Univ Paris Est Creteil and Université Paris Cité, CNRS, LISA, F-94010 Créteil, France

<sup>4</sup>Institut des Sciences Moléculaires (ISM), CNRS, Univ. Bordeaux, 351 cours de la Libération, Talence 33400, France

<sup>a)</sup> Author to whom correspondence should be addressed: [myriam.drissi@synchrotron-soleil.fr](mailto:myriam.drissi@synchrotron-soleil.fr)

<sup>b)</sup> Electronic mail: [jean-christophe.loison@cnrs.fr](mailto:jean-christophe.loison@cnrs.fr)

## ABSTRACT

Linear carbon chains are an important family of molecules detected in various astrophysical environments. In this work, we investigate the vacuum ultraviolet photoionization of HCCS, the smallest sulfur-bearing carbon chain in the  $\text{HC}_n\text{S}$  family. The radical is produced from thiirane *in situ* in a flow-tube reactor coupled to a microwave discharge. The vibronic structures are assigned using *ab initio* calculations. The adiabatic ionization energies toward the ground ( $X^+ \ ^3\Sigma^-$ ) and first electronic excited ( $a^+ \ ^1\Delta$ ) states of the cation are measured experimentally for the first time at  $9.191 \pm 0.003$  and  $9.856 \pm 0.003$  eV, respectively, while a tentative value of  $10.364 \pm 0.006$  eV is offered for the  $b^+ \ ^1\Sigma^+$  state.

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## I. INTRODUCTION

The thioketenyl radical (HCCS) is the smallest sulfur-bearing carbon chain in the  $\text{HC}_n\text{S}$  family. It has been detected in both its neutral and cationic forms in the Taurus molecular cloud (TMC-1).<sup>1,2</sup> This discovery is part of a broader trend, as numerous carbon chain molecules have been observed in the interstellar medium (ISM). These include not only pure carbon chains but also those incorporating heteroatoms such as nitrogen, oxygen, and sulfur, as well as metal-bearing elements such as magnesium and metalloid elements such as silicon. Such molecules are observed in diverse astrophysical environments, including star-forming regions, protoplanetary disks, young stellar objects (YSOs), and external galaxies, highlighting the ubiquity of this molecular family.<sup>3</sup> Among these objects, low-mass and high-mass YSOs are found to exhibit distinct chemical differentiation during their formation process, leading to regions enriched either in carbon chains or in complex organic molecules (COMs). Several factors have been proposed to drive this differentiation, including temperature, the intensity of the radiation field, and the impact of cosmic rays; however, their relative contributions remain uncertain.<sup>4–6</sup> In the case of low-mass YSOs, Spezzano *et al.*<sup>7</sup> suggested that variations in the far-ultraviolet component of the

interstellar radiation field could play a key role. This hypothesis is supported by theoretical results from two different models,<sup>8,9</sup> both indicating a potential link between far ultraviolet radiation and the observed chemical diversity.

In this context, the study of carbon chains exposed to vacuum ultraviolet (VUV) radiation is of interest for extracting key information such as ionization energies (IEs), cation fragmentation onsets, and photoionization cross sections. From a more fundamental point of view, such an interaction allows extracting spectroscopic information, which is needed to benchmark theoretical models. These data can also be used to identify and quantify reaction products in laboratory experiments,<sup>10</sup> thereby contributing to the refinement of our understanding of sulfur chemistry, where models sometimes show discrepancies with observed molecular abundances.<sup>1,2</sup>

The electronic structure of HCCS has been extensively investigated as a prototypical tetra-atomic linear molecule with a  $^2\Pi$  electronic ground state, which exhibits a Renner–Teller effect. Both experimental<sup>11–21</sup> and theoretical<sup>22–31</sup> studies have contributed to our understanding of this system. In contrast, the  $\text{HCCS}^+$  cation has only been characterized theoretically. Initial studies employed Hartree–Fock and Møller–Plesset perturbation theory, predicting a  $^3\Sigma^-$  electronic ground state for the cation.<sup>32</sup> Later, Puzzarini<sup>33</sup>

conducted a more comprehensive investigation using the CCSD(T) method. This work examined the neutral, cationic, and anionic species of HCCS, reporting adiabatic ionization energies (IEs) of 9.114 and 9.958 eV toward the ground and first electronic excited states of the cation, respectively. In addition, the study provided equilibrium geometries, dipole moments, and electron affinity values. To the best of our knowledge, no experimental data are currently available for HCCS<sup>+</sup>. In the present work, the H-abstraction reaction in a flow tube reactor with thiirane (*c*-C<sub>2</sub>H<sub>4</sub>S) is employed to produce the HCCS radical. Using VUV synchrotron radiation in the 8.9–11.0 eV region and the threshold photoelectron spectroscopy (TPES) technique, we have measured for the first time the photoelectron spectroscopy of this radical.

During the review process of this article, a study on HCCS<sup>+</sup> has been published by Michielan *et al.*,<sup>34</sup> presenting its infrared action spectrum between 450–1850 and 3000–3350 cm<sup>-1</sup>, allowing assignment of vibrational bands for both the triplet and the singlet lowest states of the cation.

## II. METHOD

### A. Experimental

Experiments were performed using a flow tube reactor coupled to a microwave (MW) discharge installed inside the permanent endstation SAPHIRS<sup>35</sup> on the undulator-based DESIRS beamline at the SOLEIL synchrotron.<sup>36</sup> The setup has been described in detail elsewhere and will only be outlined briefly here.<sup>37</sup>

The flow-tube is composed of a main tube reactor and a collinear sliding injector. Fluorine atoms are produced by MW discharge of a mixture of F<sub>2</sub> (Air Liquide, 5% in He) and He (Air Liquide, 99.995%), which is fed into the main reactor. The precursor, thiirane (*c*-C<sub>2</sub>H<sub>4</sub>S) (Sigma-Aldrich, 98%), is kept at -20 °C inside a thermostated bath. A needle valve is used to adjust the precursor density fed into the injector. The reaction time between fluorine atoms and thiirane is controlled by adjusting the position of the injector inside the reactor. The total pressure in the reactor is 6.5 × 10<sup>-1</sup> mbar for a total flow of around 1000 sccm (standard cubic centimeters per minute). The concentrations of fluorine and thiirane are estimated to be around 1 × 10<sup>13</sup> and 3 × 10<sup>13</sup> cm<sup>-3</sup>, respectively. The F atoms can react with the precursor and abstract three H atoms to generate the HCCS molecule. The experimental conditions, including precursor concentration and injector distance, are optimized to get the best signal-to-background ratio. A mass spectrum of the reactor content is provided in the [supplementary material](#).

The reactor content is skimmed twice before entering the ionization chamber, where it crosses the linearly polarized VUV synchrotron beam at a right angle in the center of the DELICIOUS III spectrometer.<sup>38</sup> A gas filter,<sup>39</sup> located upstream, is filled with krypton to ensure a high spectral purity, removing the harmonics from the undulator spectrum. The resulting electrons and ions are accelerated in opposite directions by a  $F = 88 \text{ V cm}^{-1}$  electric field toward a velocity map imaging (VMI) and a Wiley–McLaren time-of-flight (TOF) analyzer, respectively. The data are sorted into correlated pairs by means of a PEPICO acquisition scheme and allow the multiplexed recording of mass-selected photoelectron images. The extraction field used to separate the particles induces a downshift of the ionization energies of 7 meV, following the well-established

formula  $6\sqrt{F}$  (cm<sup>-1</sup>).<sup>40</sup> All values given below were shifted up accordingly.

The spectrum was recorded by scanning the photon energy between 8.9 and 11 eV with a 6 meV step and was normalized by the recorded photon flux based upon a Si photodiode measurement. The absolute energy calibration was achieved with the atomic sulfur ionization energy, the sharp sulfur autoionization features between 10.59 and 10.73 eV, and the Kr 4p<sup>5</sup> 5s (3/2) and 4p<sup>5</sup> 5s (1/2) absorption lines from the gas filter. This led to an absolute energy scale accuracy of 3 meV over the full energy range. The photon resolution was set to 0.72 Å (6 meV at 10 eV). The total (electron + photon) energy resolution was measured at 14 meV from the atomic signals.

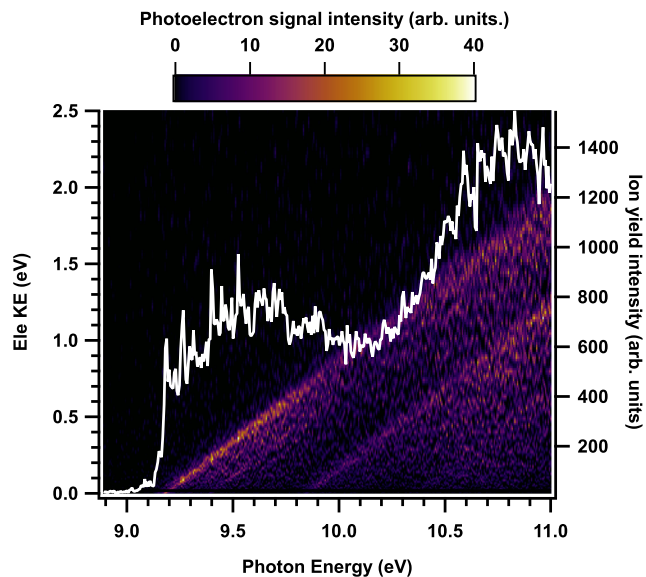
The ion image was used to select a “Region of Interest” corresponding solely to ions possessing a velocity component along the molecular beam axis, therefore removing the contribution from species not originating from the flow tube reactor and discarding contributions from the ionization chamber background (as previously described<sup>37</sup>). Furthermore, the ion image analysis provided an estimation of the translational temperature of around 300 K, which is assumed equal to the rotational temperature.<sup>41</sup>

### B. Computational

The Franck–Condon (FC) factors for photoionization were calculated using the ezSpectrum software<sup>42</sup> with geometry optimization and harmonic frequency calculations performed at the DFT level applying the M06-2X functional and the AVTZ basis set using Gaussian 2016 software.<sup>43</sup> The harmonic approximation was applied for harmonic frequencies and normal modes in both the neutral and the cationic ground states, and the Condon approximation was applied for the dipole moment. The Duschinsky effect was considered using recursive formulas already implemented in ezSpectrum. The potential energy surfaces of the two lowest electronic states of HCCS and the four lowest electronic states of HCCS<sup>+</sup> were computed as a function of the H–C–CS angle (all other parameters being optimized) using the AVTZ basis set and the MOLPRO 2015 package.<sup>44</sup> The *ab initio* calculations on the electronic states were carried out using the internally contracted multireference configuration interaction method with the Davidson correction (MRCI + Q) and complete active space self-consistent field (CASSCF) wave functions. The CASSCF and MRCI calculations were performed with 7 (6 for HCCS<sup>+</sup>) electrons distributed in 7 orbitals.

## III. RESULTS AND DISCUSSION

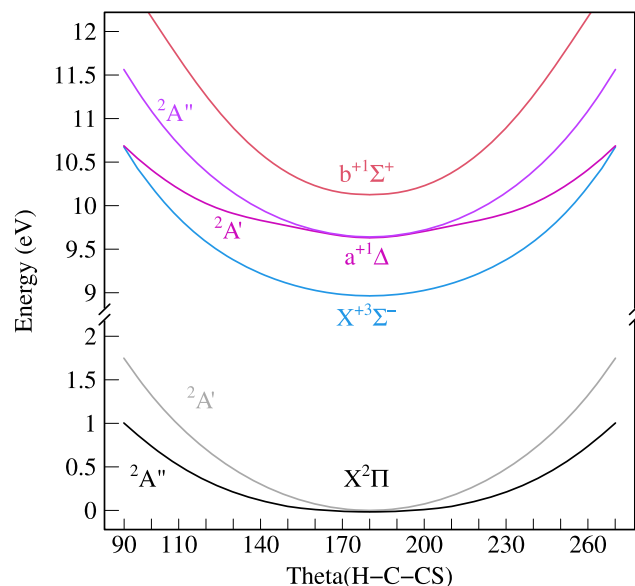
The mass-selected photoelectron spectrum (PES) matrix of HCCS (*m/z* 57) between 8.9 and 11 eV is displayed in Fig. 1. Such a matrix is built by performing the Abel transformation on the electron image to extract the electron kinetic energy for each photon energy. A wealth of information can be extracted from this 2D matrix. For instance, the ion yield can be obtained by integrating the total signal at each photon energy, as presented in white in Fig. 1. Some sharp features, assigned to autoionizing Rydberg states, can be observed on this ion yield; however, their exact assignment is beyond the scope of this article. Another way to reduce this matrix is by integrating the photoelectron signal along constant ionic state lines, a method described by Pouilly *et al.*,<sup>45</sup> yielding the so-called slow photoelectron spectrum (SPES). In this work, only slow electrons with kinetic energies up to 50 meV were considered;



**FIG. 1.** Mass-selected photoelectron signal as a function of the photon energy (horizontal axis) and the electron kinetic energy (vertical axis), for  $m/z$  57. The ion yield, obtained by integrating the signal over all electron energies, is represented on top of the matrix by a white solid line.

therefore, all resulting spectra will be referred to as threshold photoelectron spectra (TPES) throughout this article.

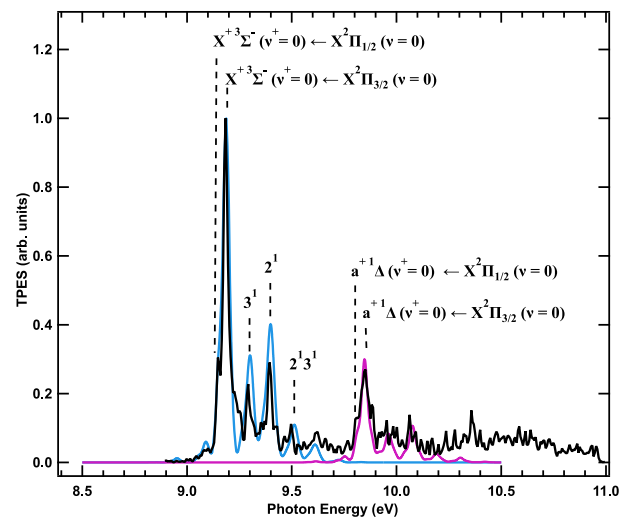
As discussed in the introduction, previous studies have shown that the  $X^2\Pi$  ground state of HCCS has a linear geometry.<sup>11</sup> Consequently, two effects will split the ground state energy: the



**FIG. 2.** Cut in the potential energy surfaces of the HCCS and HCCS<sup>+</sup> molecules as a function of the H–C–CS bending angle calculated at the MRCI+Q/AVTZ level. The other distances and angles were optimized.

Renner–Teller effect and spin–orbit coupling. In the case of the neutral ground state, the Renner–Teller effect lowers the linear symmetry to  $C_s$  and splits the state into two components,  $^2A''$  and  $^2A'$ , along the H–C–C bending coordinate (see Fig. 2). In addition, spin–orbit coupling further splits the  $^2\Pi$  state into  $\Omega = 1/2$  and  $\Omega = 3/2$  components, with the  $3/2$  component lying lower in energy ( $A_{SO} = -276.56 \text{ cm}^{-1}$ ).<sup>13</sup> Assuming the available thermal energy equals the translational temperature determined from the ion image, the population ratio between the two spin–orbit components is estimated to be 0.3. As a result, the ionizing transition is expected to originate predominantly from the  $X^2\Pi_{3/2}$  state, although we note that the above assumption of thermal equilibrium is not correct, as shown by the much higher vibrational temperature extracted from the vibrational branching ratios in the TPES. The mass-selected TPES of  $m/z$  57 is shown in Fig. 3, together with the simulated spectra whose origin has been determined to best reproduce the experimental trace. Since the HCCS ground state and HCCS<sup>+</sup> first excited electronic state are affected by the Renner–Teller effect, a FC simulation may be insufficient to fully reproduce the observed spectral structure due to the breakdown of the Born–Oppenheimer approximation. Nevertheless, in this case, the simple FC calculation with a vibrational temperature of 1000 K already shows good agreement with the experimental spectrum.

Comparison with the *ab initio* calculation allows the identification of two origin bands: the first at  $9.191 \pm 0.003 \text{ eV}$  and the second at  $9.856 \pm 0.003 \text{ eV}$ . These, together with the computed values, are summarized in Table I. Each of the origin bands presents a shoulder on the left side, which is well reproduced in the simulation



**FIG. 3.** Mass-selected experimental TPES of HCCS in the vicinity of  $X^+3\Sigma^- \leftarrow X^2\Pi$  and  $a^+1\Delta \leftarrow X^2\Pi$  ionizing transitions (black trace) compared to Franck–Condon simulations performed with  $T_{\text{vib}} = 1000 \text{ K}$  (blue and pink traces, respectively). The relative intensities of the simulations have been adjusted to reproduce the experimental trace. The origins of the FC simulations have been determined as the best match to our experimental TPES. The vibrational assignments follow the usual notation, where the mode index is given by the main number and the superscript indicates the number of quanta excited in that mode, and they all correspond to the  $X^+3\Sigma^- \leftarrow X^2\Pi_{3/2}$  band.

**TABLE I.** Experimental and calculated ionization energies in eV of HCCS in the vicinity of  $X^+ {}^3\Sigma^- \leftarrow X^2\Pi_{3/2}$ ,  $a^+ {}^1\Delta \leftarrow X^2\Pi_{3/2}$  and  $b^+ {}^1\Sigma^+ \leftarrow X^2\Pi_{3/2}$  ionizing transitions. Note that the spin-orbit splitting is not taken into account in the calculated values.

	Experimental	M06-2X/AVTZ	MRCI+Q/CBS	CCSD(T) <sup>33</sup>
$X^+ {}^3\Sigma^- \leftarrow X^2\Pi_{3/2}$	9.191 ± 0.003	8.90	9.00	9.114
$a^+ {}^1\Delta \leftarrow X^2\Pi_{3/2}$	9.856 ± 0.003	10.21	9.68	9.958
$b^+ {}^1\Sigma^+ \leftarrow X^2\Pi_{3/2}$	...	...	10.16	...

using the SO constant derived by He and Clouthier ( $A_{SO} = -276.56 \text{ cm}^{-1} = -34.3 \text{ meV}$ )<sup>13</sup> and our estimated SO-component population ratio. The relative ionization efficiency of each photoionization transition can be estimated as detailed in our prior work.<sup>46</sup> Briefly, the relative photoionization efficiency for each transition is estimated considering only transitions involving the ejection of a single electron and assuming similar electronic wavefunction overlaps. In the case of HCCS, the neutral electronic ground state has an electronic configuration equal to  $(1-9)\sigma^2(1-2)\pi^43\pi^3$ , and the three lowest electronic states of HCCS<sup>+</sup>, accessible by direct ionization, are all issued from the  $(1-9)\sigma^2(1-2)\pi^43\pi^2$  electronic configuration. The probability of the photoionizing transition is, in our simplified approximation, proportional to the weighted product of the initial and final main configurations involved in single electron ionization processes, times the degeneracy of the projection of the orbital angular momenta and the spin multiplicities of the initial and final electronic states. The relative photoionization intensities from HCCS( $X^2\Pi$ ) toward HCCS<sup>+</sup> in the  $X^+ {}^3\Sigma^-/a^+ {}^1\Delta/b^+ {}^1\Sigma^+$  states are then equal to 3/2/1. These theoretical intensities are in qualitative agreement with the observed intensities. However, it should be noted that photoionization to the HCCS<sup>+</sup> ( $X^+ {}^3\Sigma^-$ ) ground state is clearly favored, likely due to the presence of autoionization resonances, which are known to affect the relative intensities of the TPES peaks.<sup>47</sup> Indeed, these resonances are clearly seen in the total ion yield spectra shown in Fig. 1.

The measured adiabatic ionization energy (AIE) of  $9.191 \pm 0.003 \text{ eV}$  can be assigned to the  $X^+ {}^3\Sigma^- \leftarrow X^2\Pi_{3/2}$  ionizing transition and is in reasonable agreement with the computed value of  $9.114 \text{ eV}$  reported by Puzzarini<sup>33</sup> and with our calculations ( $8.90 \text{ eV}$  at the M06-2X/AVTZ level and  $9.00 \text{ eV}$  at MRCI+Q/CBS with geometry and ZPE calculated at the AVTZ level). According to our calculations, the vibrational structure in the TPES arises mainly from the  $\nu_2^+$  and  $\nu_3^+$  modes, corresponding to the C–C and C–S stretching modes, respectively. The observed peaks are assigned to the adiabatic transition, the  $\nu_2^+$  and  $\nu_3^+$  fundamentals, and the  $\nu_2^+ + \nu_3^+$  combination band. The observed frequencies for the  $\nu_2^+$  and  $\nu_3^+$  modes,  $1701 \pm 48$  and  $871 \pm 48 \text{ cm}^{-1}$ , respectively, show good agreement with the work of Michielan *et al.*<sup>34</sup> In particular, the  $\nu_2^+$  frequency is in agreement with their experimental value, and both modes are consistent with the theoretical frequencies reported therein. Note that the mode numbering used here follows the notation of He and Clouthier<sup>11</sup> and differs from that adopted by Michielan *et al.*<sup>34</sup>

The second electronic band corresponds to the  $a^+ {}^1\Delta \leftarrow X^2\Pi_{3/2}$  ionizing transition, where the final state is expected to undergo Renner–Teller splitting into two components, as shown in Fig. 2. Here, both states are linear with a small geometry change. However,

the Renner–Teller effect will probably affect the vibrational structure along the RT active modes ( $\nu_4$  and  $\nu_5$  bending modes), although we have not observed any RT signature in our TPES. In this band, due to the lower signal-to-background from the abovementioned statistical arguments, only the adiabatic transition could be fitted, resulting in an AIE of  $9.856 \pm 0.003 \text{ eV}$ . This value is significantly lower than the value calculated by Puzzarini<sup>33</sup> for the first singlet electronic state of HCCS<sup>+</sup>,  $9.958 \text{ eV}$ , labeled the  ${}^1\Sigma^+$  state in their paper, probably due to an unsuitable choice for the reference determinant or the multi-reference nature of this molecule. The experimental value is also significantly lower than calculated at the M06-2X/AVTZ level, equal to  $10.21 \text{ eV}$ . This theoretical overestimation of the IE is likely due to the multi-configurational aspect of the  $a^+ {}^1\Delta$  electronic state of HCCS<sup>+</sup>. The multiconfiguration MRCI+Q/CBS method gives slightly better results,  $9.68 \text{ eV}$ , but underestimates the IE value, as in the case of ionization to the ground state of HCCS<sup>+</sup>. Finally, we note that the recent experimental study by Michielan *et al.*<sup>34</sup> reports an energy separation of  $0.23 \text{ eV}$  ( $22 \text{ kJ/mol}$ ) between the ground state and the first excited state of HCCS<sup>+</sup>, which is inconsistent with the TPES measurements ( $0.665 \text{ eV}$ ) and the calculation ( $0.68 \text{ eV}$ ) presented here.

The weak signal peaking at  $10.364 \text{ eV}$  in the TPES may be due to the photoionization toward the second excited state  $b^+ {}^1\Sigma^+ \leftarrow X^2\Pi$  calculated at  $10.16 \text{ eV}$  at the MRCI+Q/CBS level, thus with a similar red shift of around  $-0.2 \text{ eV}$  to the MRCI+Q/CBS calculations for  $X^+ {}^3\Sigma^- \leftarrow X^2\Pi_{3/2}$  and  $a^+ {}^1\Delta \leftarrow X^2\Pi_{3/2}$  ionizing transitions. In addition, its relative intensity is consistent with our statistical estimate, i.e., twice as weak as the  $a^+ {}^1\Delta \leftarrow X^2\Pi$  ionizing transition. We, therefore, tentatively associate an adiabatic ionization energy of  $10.364 \pm 0.006 \text{ eV}$  to the  $b^+ {}^1\Sigma^+$  state, but we note that the low signal-to-background precludes a categorical assignment. For this reason, this AIE is not included in Table I.

#### IV. CONCLUSIONS

The thioketenyl radical photoelectron spectrum has been recorded in the region of the ground and first excited states for the first time using a double imaging photoelectron photoion spectrometer. HCCS was produced through H abstractions of thiirane with fluorine atoms in a discharge flow-tube reactor. The two intense vibronic bands were assigned with the aid of *ab initio* calculations, yielding the adiabatic ionization energies. The first excited state was determined to be  ${}^1\Delta$ , in contrast to earlier calculations that assigned it as  ${}^1\Sigma^+$ . This work is the first experimental determination of the adiabatic ionization energies of HCCS. These data provide a foundation for use in thermochemical calculations, such as determining

the enthalpy of formation or the dissociation energy of the cation.<sup>48</sup> However, since the corresponding thermochemical parameters of the neutral species remain unknown, their determination is necessary before these derived quantities can be accurately evaluated. In addition, the photoelectron spectrum acts as a fingerprint that can be used to detect HCCS in advanced mass spectrometry experiments in gas phase reaction studies.<sup>10</sup> Finally, it is worth noting that experimental values are still needed to benchmark theoretical methods, where even for relatively small systems, what many consider the gold standard [CCSD(T)] fails, and multireference/multiconfigurational methods are off by 0.2 eV.

## SUPPLEMENTARY MATERIAL

See the [supplementary material](#) for the TOF mass spectrum of the reactor content.

## ACKNOWLEDGMENTS

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## AUTHOR DECLARATIONS

### Conflict of Interest

The authors have no conflicts to disclose.

## Author Contributions

**M. Drissi:** Data curation (lead); Formal analysis (lead); Investigation (equal); Writing – original draft (lead); Writing – review & editing (equal). **G. A. Garcia:** Formal analysis (supporting); Investigation (equal); Methodology (equal); Resources (equal); Supervision (lead); Writing – review & editing (equal). **B. Gans:** Formal analysis (equal); Investigation (equal); Writing – review & editing (equal). **S. Boyé-Péronne:** Investigation (equal); Writing – review & editing (equal). **H. R. Hrodmarsson:** Investigation (equal); Writing – review & editing (equal). **L. Nahon:** Investigation (supporting); Methodology (supporting); Supervision (equal); Writing – review & editing (equal). **J.-C. Loison:** Formal analysis (equal); Investigation (equal); Methodology (equal); Supervision (equal); Writing – review & editing (equal).

## DATA AVAILABILITY

Raw data were generated at the SOLEIL Synchrotron large scale facility. Derived data supporting the findings of this study are available from the corresponding author upon reasonable request.

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